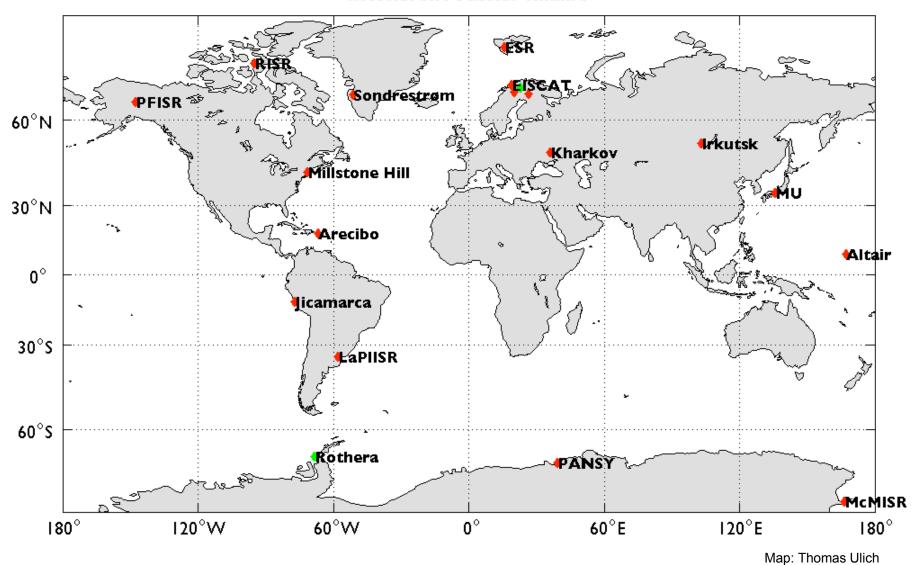
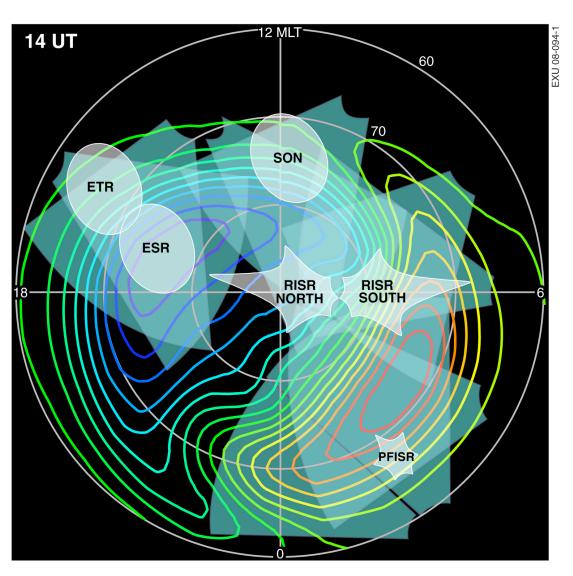
# The US Incoherent Scatter Radars

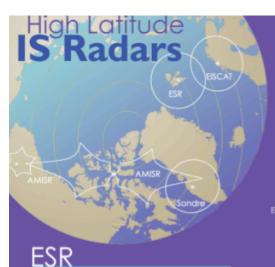
Anja Strømme
SRI International
With contributions from the other
PIs

#### **Incoherent Scatter Radars**



# Map of the north...





#### Useful Constants

c<sub>n</sub> = 2.99792458 x 10<sup>8</sup> ms<sup>-1</sup>  $\mu_0 = 4\pi \times 10^{-7} \text{ NA}^{-2}$ 1u = 1.66053873 x 10-27 kg  $G = 6.673 \times 10^{-11} \text{ m}^3 \text{kg}^{-1} \text{s}^{-1}$ 

EISCAT2003

Plasma frequency,  $\omega_D = \left(\frac{n_e e^2}{r_e me}\right)^{1/2} = 2\pi \sqrt{80.6 n_e}$ Debye length,  $\lambda_0 = \left(\frac{\epsilon_0 k_B T_e}{n_e e^2}\right)^{1/2} = 69 \sqrt{\frac{T_e}{n_e}}$ 

Ion gyro frequency,  $\Omega_i = \frac{-eB}{mi} = 9575 \frac{B}{W}$ 

W: molecular weight; m = W • u

67°21'49" N

26° 37'37" E 76° 43'

Fully steerable

UT + 2 (UT + 3 summer)

63° 34'N

Geomag. dip angle Invariant latitude Local time

Magnetic time

Elevation

Geomag, dip angle Invariant latitude

69° 35'11" N 19° 13'38' E 77° 30' 66°12' N

78° 09"11" N

75° 10'48" N

360° az., above, 30° elev.\*

16° 1'44" E

82° 06'

VHF: ±15° az., 30°-90° elev. (north only)

77° 32' Geomag, dip angle 66° 7' N Local time Magnetic time

65° 7'12" N 74° 43' 46" N. 147° 25' 48" W 94° 54' 16" E

88° 47° UT - 6 (central) UT - 7:45

#### Sondrestrom

67° 51'38" N

Fully steerable

64° 27' N

309° 03'02" E Geomag. dip angle 80° 24' Invariant latitude Local time Magnetic time 74° 11'24" N

UT - 3 (UT - 2 summer) UT - 1:58 360° az., above 25°-30° elev.\*

## The Incoherent Scatter RADAR Equation: $Signal, P_r = \frac{Cc_0G\lambda^2}{2} \frac{P_1\tau_P}{r^2} \frac{\sigma_e n_e(r)}{(1+k\lambda_D^2)(1+k\lambda_D^2+T_r)}$

Center transmit frequency, f 500 MHz

10.48 m<sup>-1</sup> (2k = 20.96 m<sup>-1</sup>)

Noise, Pn = PnknTsvs BW : BW - bandwidth

 $\sigma_0 = 4\pi r_0^2 = 9.98 \times 10^{-29} \text{ m}$  . Radar cross section of

0.5996 m (\(\lambda\)/2 = 0.2998 m) C = Constant containing loss and antenna shape T. MIW 25% <1 - 2000 us 32 m / 42 m 42.5 dBi / 45 dBi 0.60 55-65 K Parabolic dish Cassegrain

Wavenumber, k

Wavelength, \(\lambda\)

Peak power, P. Max duty cycle

Pulse length, To

Antenna size (dia.)

Antenna gain, G

Polarization

Antenna beamwidth

System temperature, Toy Antenna type

Center transmit frequency, f Wavenumber, k Wavelength, λ Peak power, Pe Max duty cycle Pulse length, To Antenna size (dia.) Antenna gain, G

Antenna beamwidth System temperature, T<sub>sys</sub> Antenna type Feed system Polarization

EISCAT VHF Center transmit frequency, f Wavenumber, k Wavelength, λ Peak power, Pr

Max duty cycle Pulse length,  $\tau_0$ Antenna size Antenna gain, G

Antenna beamwidth\* System temperature, T<sub>sys</sub> Antenna type Feed system Polarization

928.4 MHz 0.3229 m (\(\lambda/2 = 0.1615 m\) 220 km\*\*, \(\lambda/2 = 0.17 m\) 10%

1-2000 us 32 m 48 dBi 0.5° 70 - 80 K Parabolic dish Cassegrain Circular

224 MHz

4.695 m<sup>-1</sup> (2k = 9.389 m<sup>-1</sup>)

32 m

0.5°

Anv

48 dBi

30 - 35 K

Parabolic dish

Cassegrain

1.3384 m (\(\lambda\)/2 = 0.6692 m) 2 x 1.5 MW 12.5% 120 x 40 m (4 x (30 x 40 m)) 46 dBi

250 - 350 K Offset parabolic cylinder Line feed Circular

449 MHz Center transmit frequency, f 9.4 m<sup>-1</sup> (2k = 18.82 m<sup>-1</sup>) Wavenumber, k Wavelength, λ 0.6677 m (\(\lambda\)/2 = 0.3338 m) Peak power, P. 2 MW 10% Max duty cycle Pulse length, To  $1 - 2000 \, \mu s$ 128 x (1.5 x 3.5 m) Antenna size Antenna gain, G 43 dBi Antenna beamwidth\* System temperature, Tour 120 K Crossed dipole phased array Antenna type

Distributed amplifiers Feed system Circular Polarization

\* Full width, half power \*\* Tromsø field line

#### Kiruna UHF Sodankylä UHF

19.46 m<sup>-1</sup> /2k = 38.92 m<sup>-1</sup> | 220 km\*\*, 2k sin 0 = 36 m<sup>-1</sup> 220 km\*\*, 2k sin 8 = 33 m-1 220 km\*\*, λ/2 = 0.19m

32 m 48 dBi 0.50 30 - 35 K Parabolic dish Cassegrain Any

Whole antenna: 0.6° EW x 1.7° NS; Half antenna: 1.2° EW x 1.7° NS

#### Sondrestrom

Wavenumber, k Wavelength, λ. Peak nower P. Max duty cycle Pulse length, To Antenna size (dia.) Antenna gain, G Antenna beamwidth\* System temperature, Trus Antenna type Feed system Polarization

27.04 m<sup>-1</sup> (2k = 54.07 m<sup>-1</sup>) 0.2323 m (\(\lambda/2=0.1162 m\) 3.5 MW 3%

 $2 - 500 \, \mu s$ 32 m 49 dB 0.5° 85 K Parabolic dish Cassegrain Circular



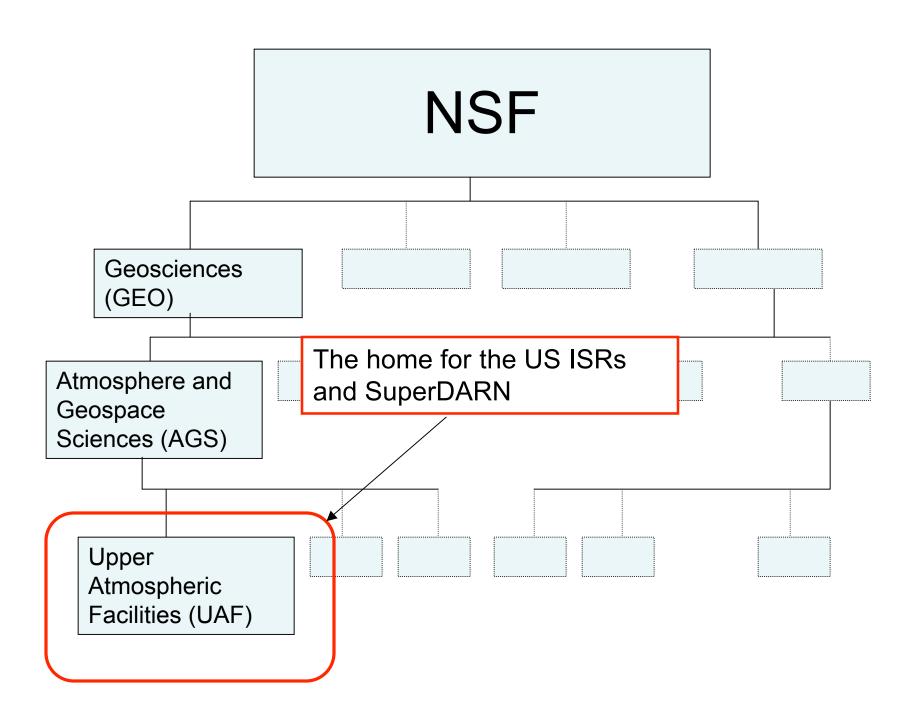


<sup>\*</sup> nitial locations. Poker Flat (one face) operational Spring 2005. Resolute Bay (two faces) operational Winter 2006. One face = 128 panels.

<sup>\*</sup>Dependent on azimuth. No transmission below this elevation.

# How it works - political

 All the US Incoherent Scatter Radars are funded by the National Science Foundation (NSF) and are governed under the Upper Atmosphere Facility (UAF) program.



# How it works - political

- All the US Incoherent Scatter Radars are funded by the National Science Foundation (NSF) and are governed under the Upper Atmosphere Facility (UAF) program.
- The Operation and Maintenance (O&M) is done by a host institution through a cooperative agreement renewed every 5 years.

## From South to North...

- Jicamarca Cornell University (in collaboration with Instituto Geofísico del Perú)
- Arecibo Cornell University (currently in the process of an open re-bid)
- Millstone Hill MIT
- PFISR SRI International
- Sondrestrom SRI International
- RISR-N SRI International

# How it works - practically (1)

- The radars are available to all US users, and in reality to all users.
- Analyzed data is available in the madrigal database
  - http://jro.igp.gob.pe/madrigal/ (Jicamarca)
  - <u>http://madrigal.naic.edu/</u> (Arecibo)
  - http://madrigal.haystack.mit.edu/madrigal/ (Millstone Hill)
  - <u>http://isr.sri.com/madrigal/</u> (SRI International)
  - <a href="http://www.eiscat.se/madrigal/">http://www.eiscat.se/madrigal/</a> (EISCAT)

## Contact information Principal Investigator

- Jicamarca: Jorge L. (Koki) Chau jorge.chau@jro.igp.gob.pe
- Arecibo: Sixto Gonzales sixto@naic.edu
- Millstone Hill: Phil Erickson pje@haystack.mit.edu
- PFISR: Mike Nicolls <u>michael.nicolls@sri.com</u>
- RISR (+AMISR): Craig Heinselman craig.heinselman@sri.com
- Sondrestrom: Anja Strømme anja.stromme@sri.com

# How it works - practically (2)

- There are annual "world days" coordinated between all the radars (653 hours in 2010).
- Otherwise scheduling is done individually for each radar (the SRI radars - Sondrestrom, PFISR and RISR - are scheduled together).
- We are working toward more coordinated scheduling and operation of all the UAF radars.

# 6 slide pre-AMISR history of the US radars (Bob Robinson):

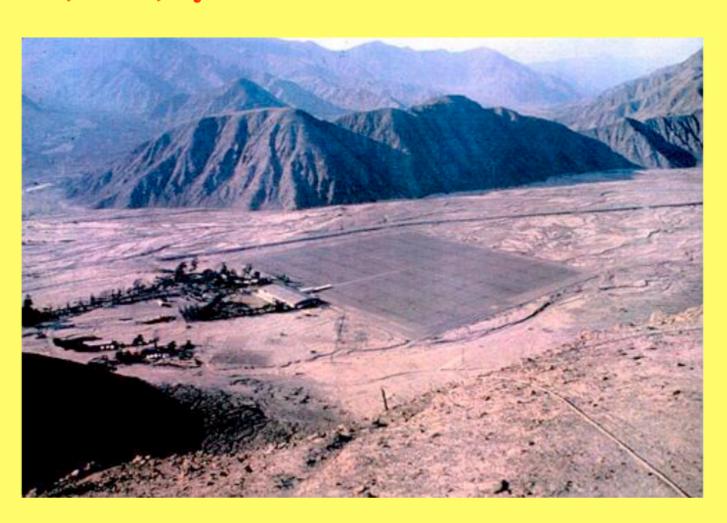


## 1958-1959

**Bill Gordon** conceives of the idea of incoherent scatter and construction begins at Arecibo, Puerto Rico, with funding from the **Defense Advanced Research Project** Office (DARPA)

## <u>1961</u>

# The Jicamarca Observatory is constructed near Lima, Peru, by the National Bureau of Standards.

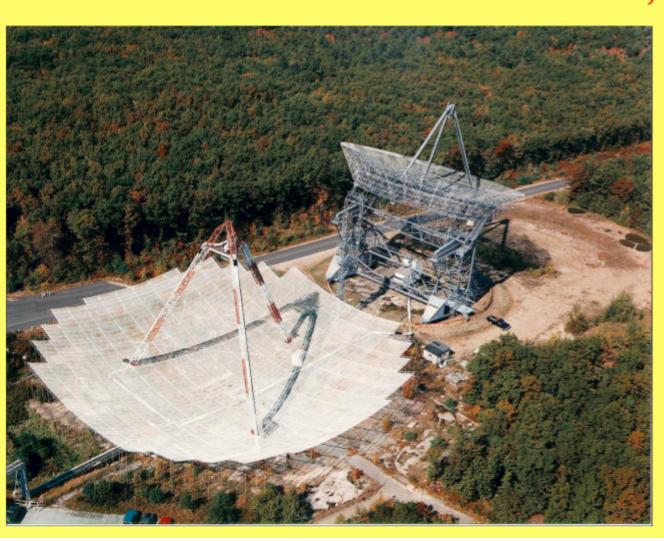


1962
Construction of Arecibo is completed.



## **1963**

The Millstone Hill zenith antenna is constructed by MIT Lincoln Laboratories at a site near Boston, MA



### <u> 1971</u>

The Chatanika Radar is moved from Stanford University to a new site near Fairbanks, Alaska.

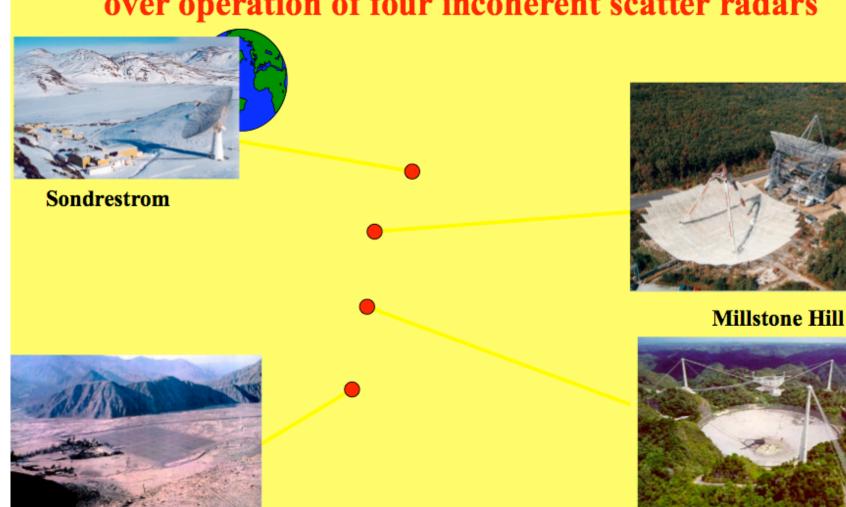




1982
The Chatanika
Radar is moved to a
new location near
Sondrestrom,
Greenland,

## <u>1970 – 1980</u>

## The U.S. National Science Foundation takes over operation of four incoherent scatter radars



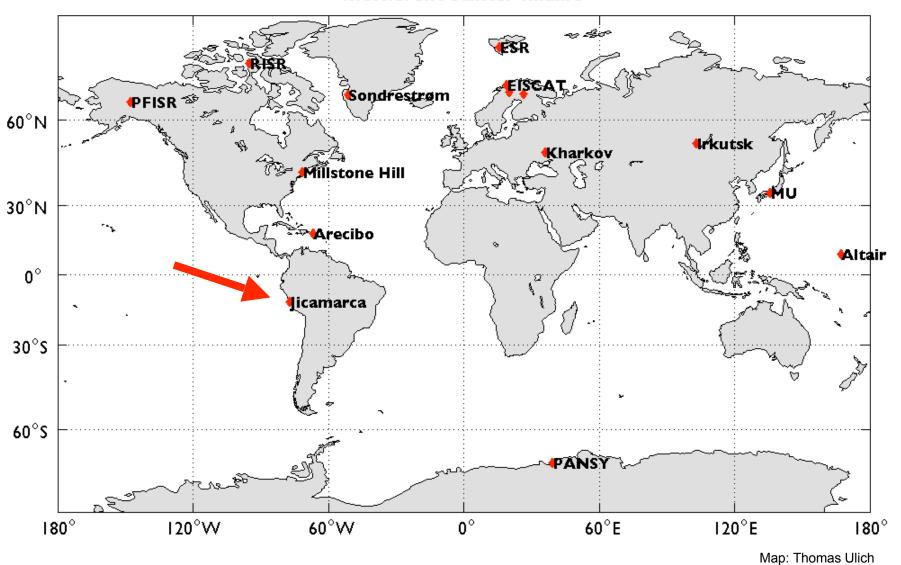
**Jicamarca** 

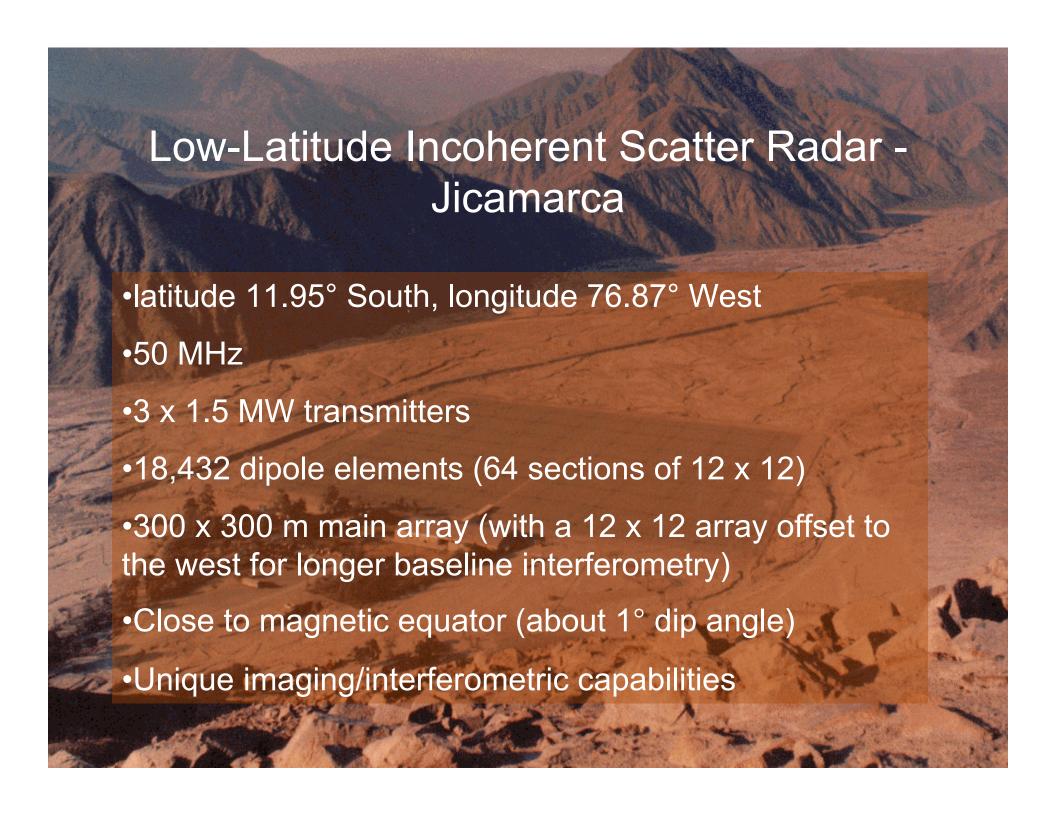


Arecibo

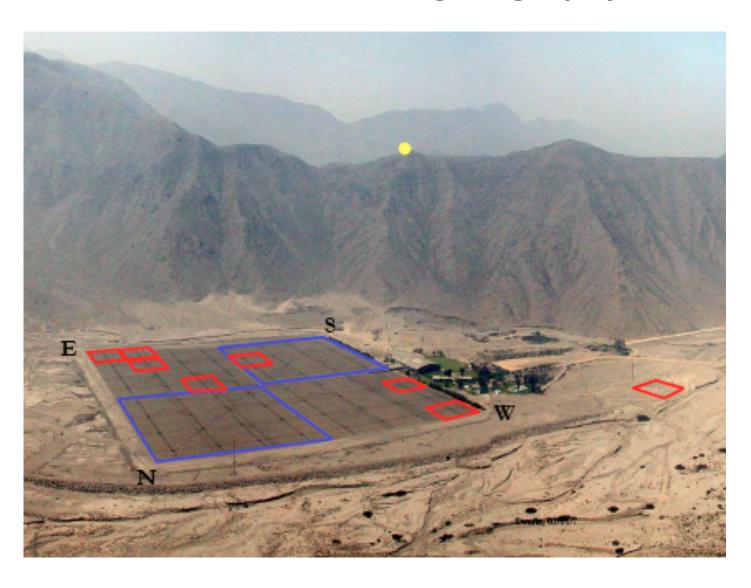
## The radars:

#### **Incoherent Scatter Radars**

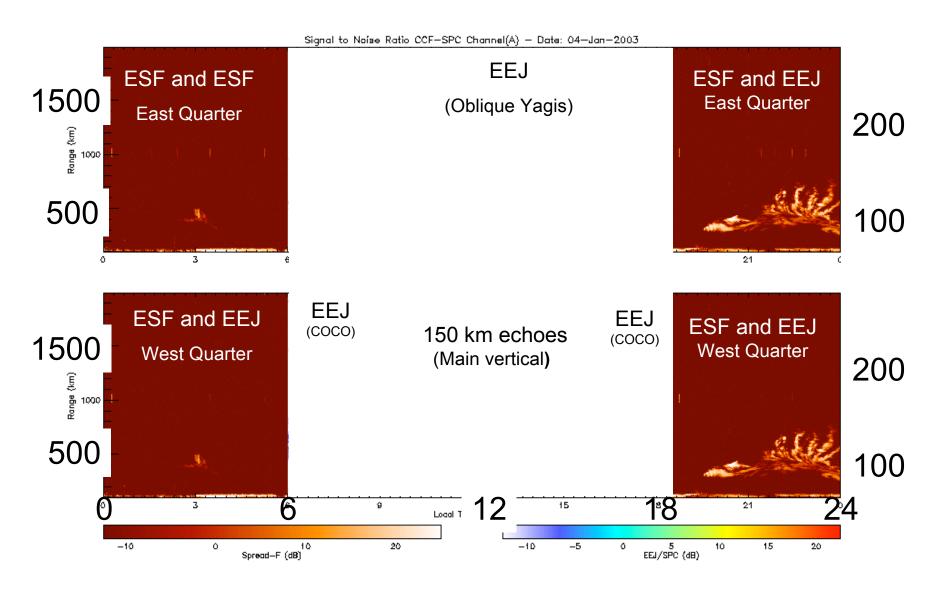




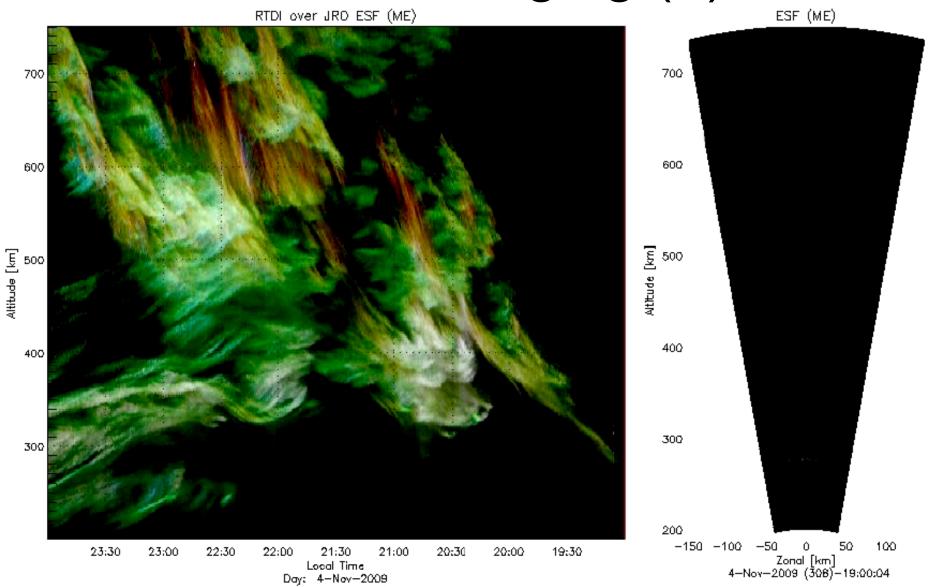
# JULIA – Imaging (1)



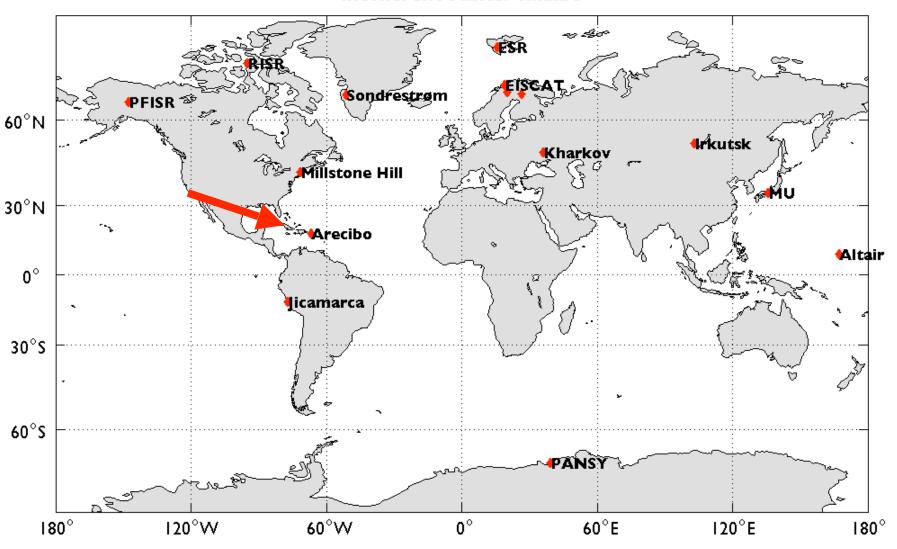
## **JULIA Observations**



# JULIA – Imaging (2)

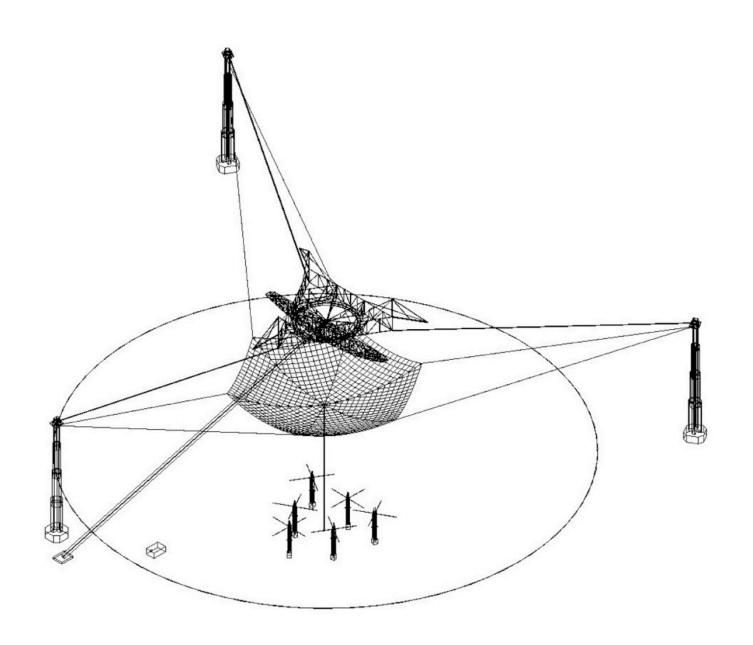


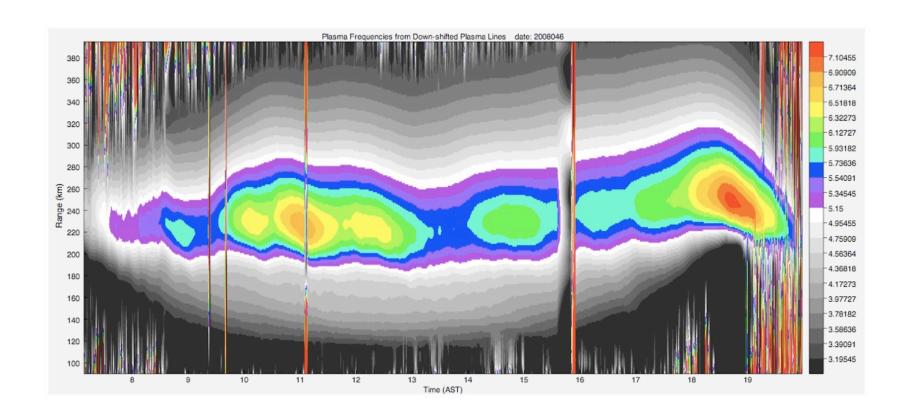
#### **Incoherent Scatter Radars**



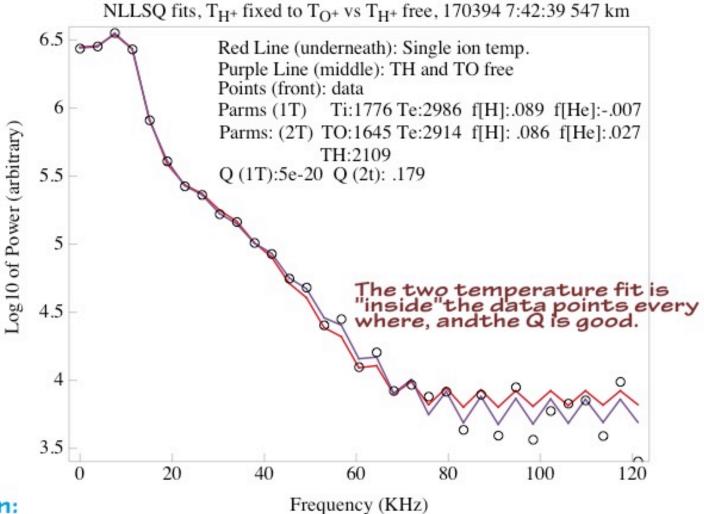
Map: Thomas Ulich

... Photo courtesy of the NAIC - Arecibo Observatory, a facility of the NSF ... Photo by David Parker / Science Photo Library Arecibo ISR •latitude 18 ° 20'36.6"North •longitude 66 ° 45'11.1"W West •430 Mhz •2.5 MW transmitters •305m diameter fixed dish Two feeds - line feed and Gregorian feed Also used as 2.4 GHz planetary radar Also used as (the worlds largest single dish) radio telescope New heater in 2011





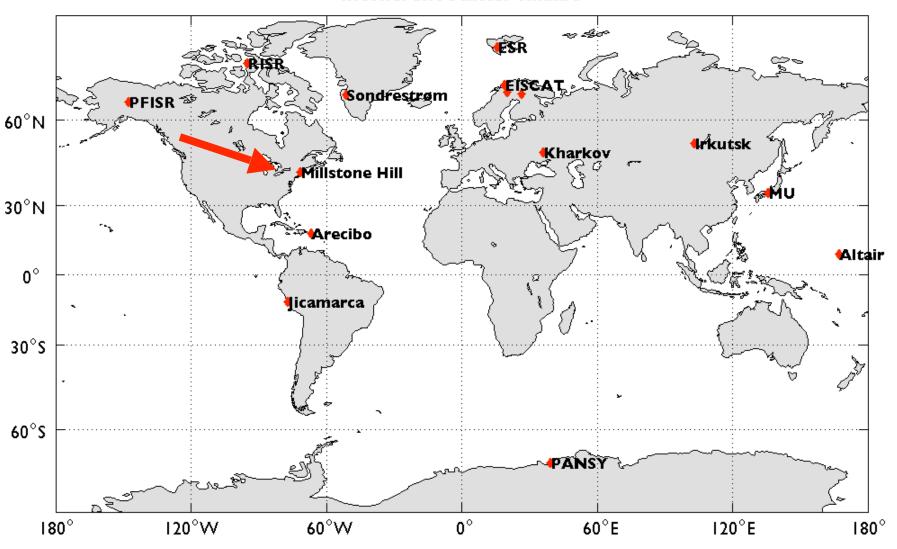
#### A Model with More Freedom



Conclusion:

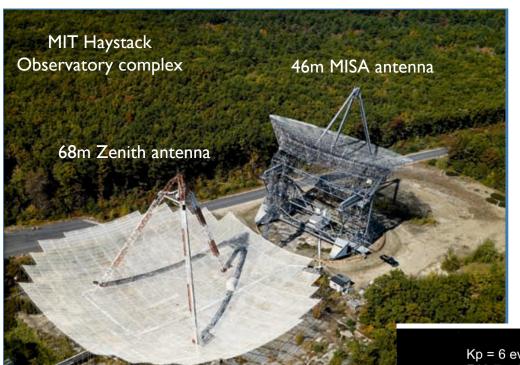
Separate 0+ and H+ temperatures are necessary and sufficient to obtain a good fit.

#### **Incoherent Scatter Radars**



Map: Thomas Ulich





Millstone Hill UHF Incoherent Scatter Radar Westford, MA USA

440.2 MHz frequency 2.5 MW peak TX System temperature ~150-200 K overall

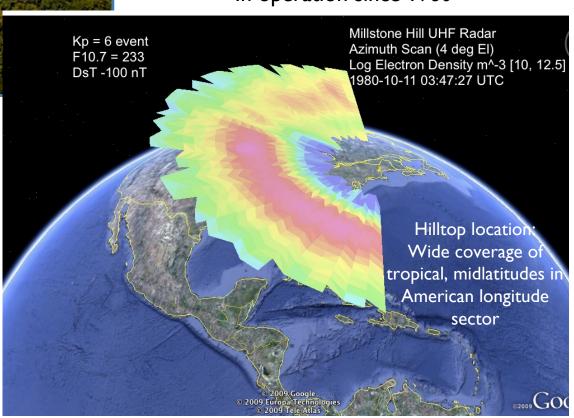
Fully steerable 46m antenna 68m zenith antenna

In operation since 1960



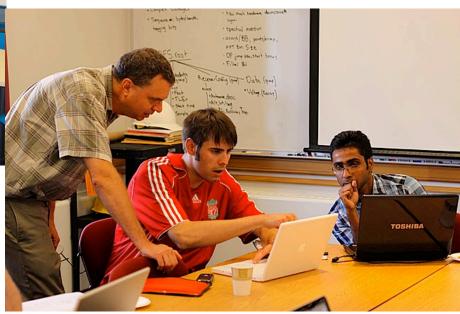
Litton 5773 UHF klystrons

2 x 1.25 MW peak 6% max duty cycle



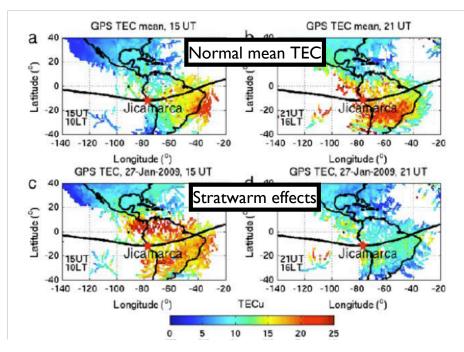


## Millstone Hill Geospace Science Center

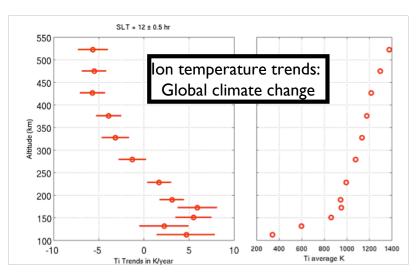


- Modernized control, receiver area
- Displays for multiple instruments
- Smartboards for interactive teaching
- Community resource for coordinated campaigns, workshops
- Focus for North American Regional Distributed Arrays (DASI) activities

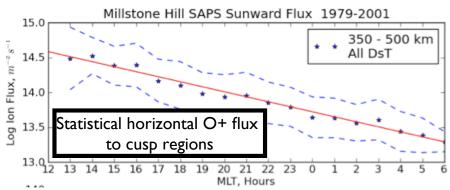
## Millstone Hill research: Geospace System Science



Lower/upper atmospheric coupling
[Stratospheric sudden warming]
and mesoscale effects on ionosphere, thermosphere
Goncharenko, Coster

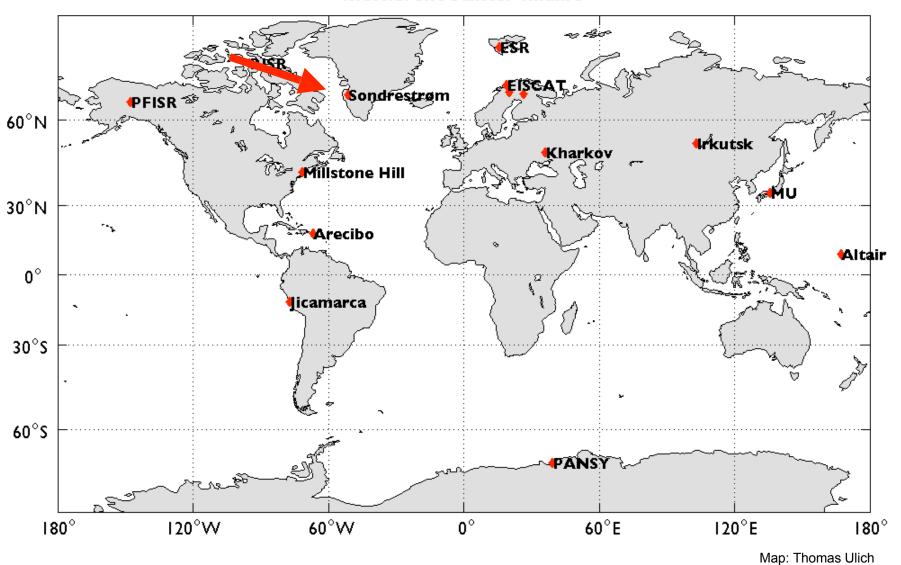


Long term ionospheric trend studies Zhang, Holt



Subauroral ionospheric redistribution in SED/SAPS regions
Erickson, Foster, Coster

#### **Incoherent Scatter Radars**



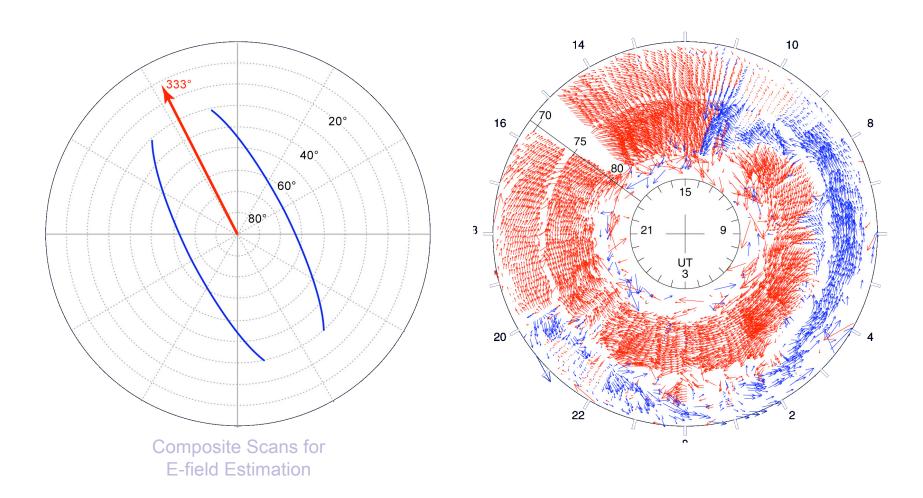


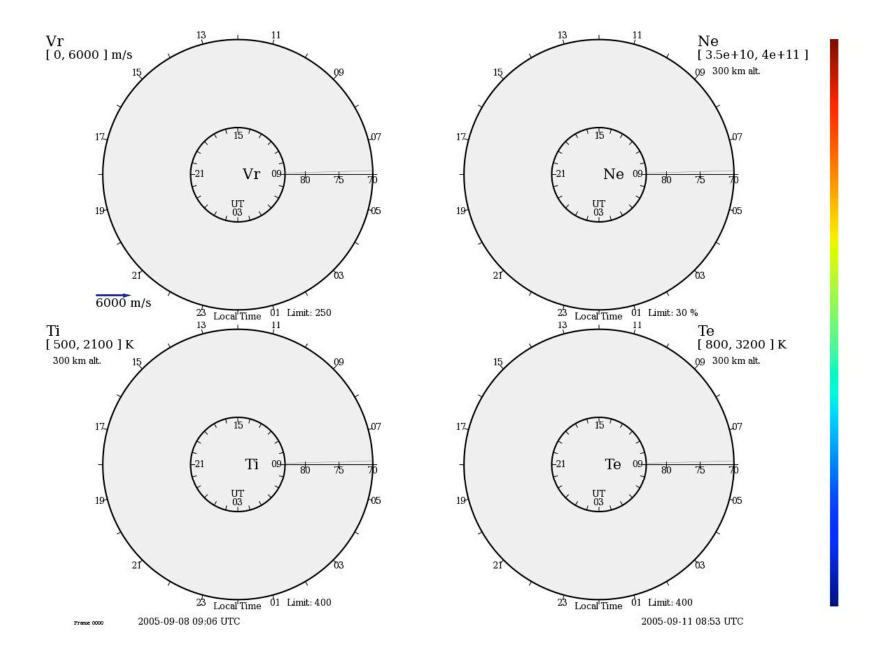
- •latitude 66 °59'12"North
- •Longitude 309° 03'02" East
- •1290 Mhz
- •3 MW transmitters
- •32m diameter steerable dish
- Colocated lidars and ASIs
- About 20 other colocated instruments
- Operational in Kangerlussuaq since 1982 (after a solar cycle in Chatanika, Alaska)



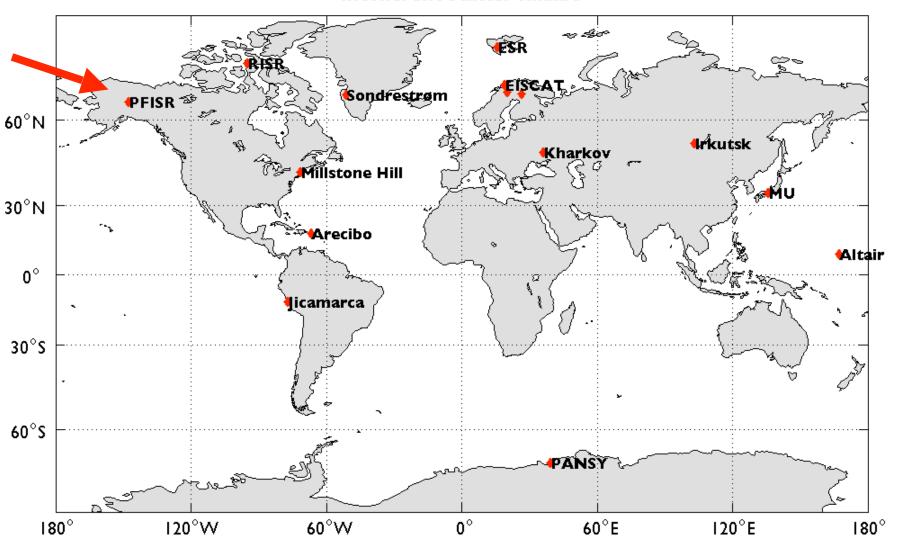


## Ion Velocity (E-field) Maps





#### **Incoherent Scatter Radars**



Map: Thomas Ulich

# ...but first some general words about AMISR...

### **AMISR**

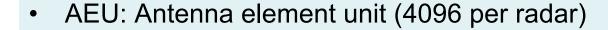
- First US ISR designed under NSF funding for pure scientific research
- First modular pulse-to-pulse steerable Incoherent scatter radar

### Abbriviations!!!!

- AMISR Advanced Modular Incoherent Scatter Radar
  - Refers to the technology and the overall kind of radar
- PFISR Poker Flat Incoherent Scatter Radar
  - An AMISR radar located in Poker Flat, Alaska
- RISR-N Resolute Bay Incoherent Scatter Radar -North (toward ESR!)
  - An AMISR Radar in Resolute Bay, Canada. Pointing northward
- RISR-S/RISR-C Resolute Bay Incoherent Scatter Radar - South/Canada
  - An AMISR radar under construction in Resolute Bay,
     Canada. Pointing South. Funded (and owned) by Canada.
     Official name RISR-C

## The lego set

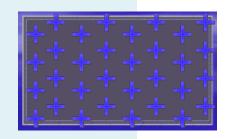
AMISR: Advanced Modular Incoherent Scatter Radar



- Panel: smallest "lego" piece (consist of 32 AEU)
- Group: Set of Panels (consist of 8 panels)
- Face: Set of groups (one complete radar ("Face") consist of 16 groups, 128 panels, 4096 AEU...)

Panel





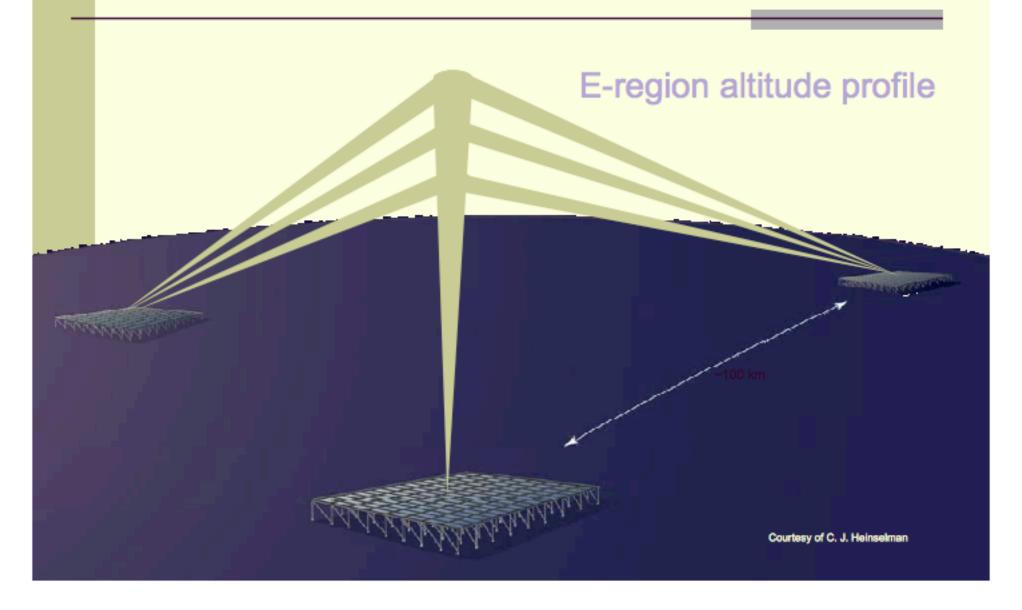
Face

### More about the "current" AMISR

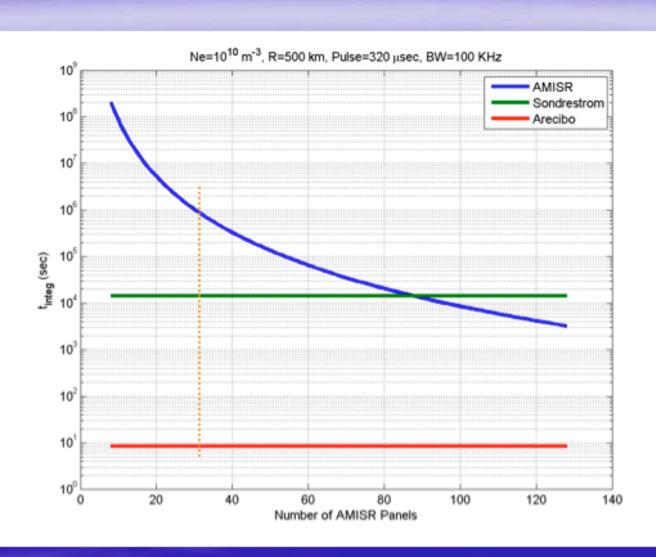
- NSF originally funded 3 full faces. Funding was sufficient to two (and a bit)
  - One operational since January 2007 in Poker Flat, Alaska (PFISR)
  - Second operational since Dec 2009 in Resolute Bay, Canada (RISR-N)
  - A collaboration with Canada will lead to the completion of the third face of the original AMISR plan (RISR-S now RISR-C).
- Modular/Transportable/Reconfigurable
- Phased array pulse-to-pulse steering
- Solid state
  - No moving parts, can hence be remotelly controlled
- Gentle degrade
- 430-450 MHz TX frequency
- ~2 MW peak power per radar (10% duty cycle)
- 1 μs to 2 ms pulses

### Tristatic Configuration

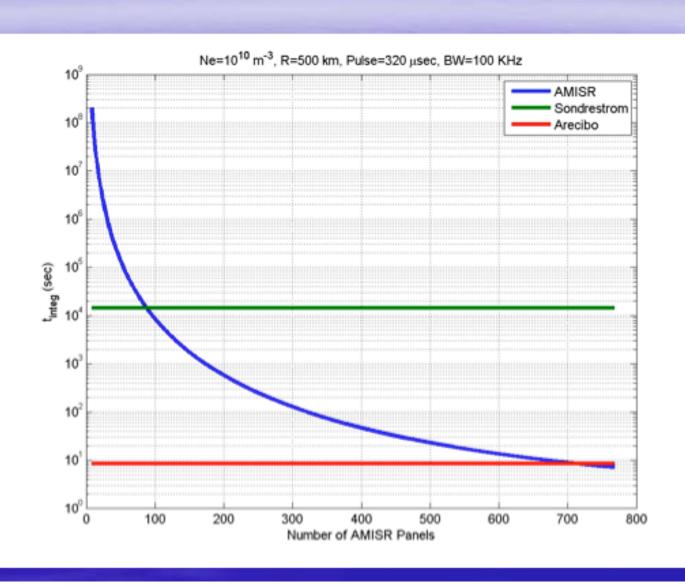


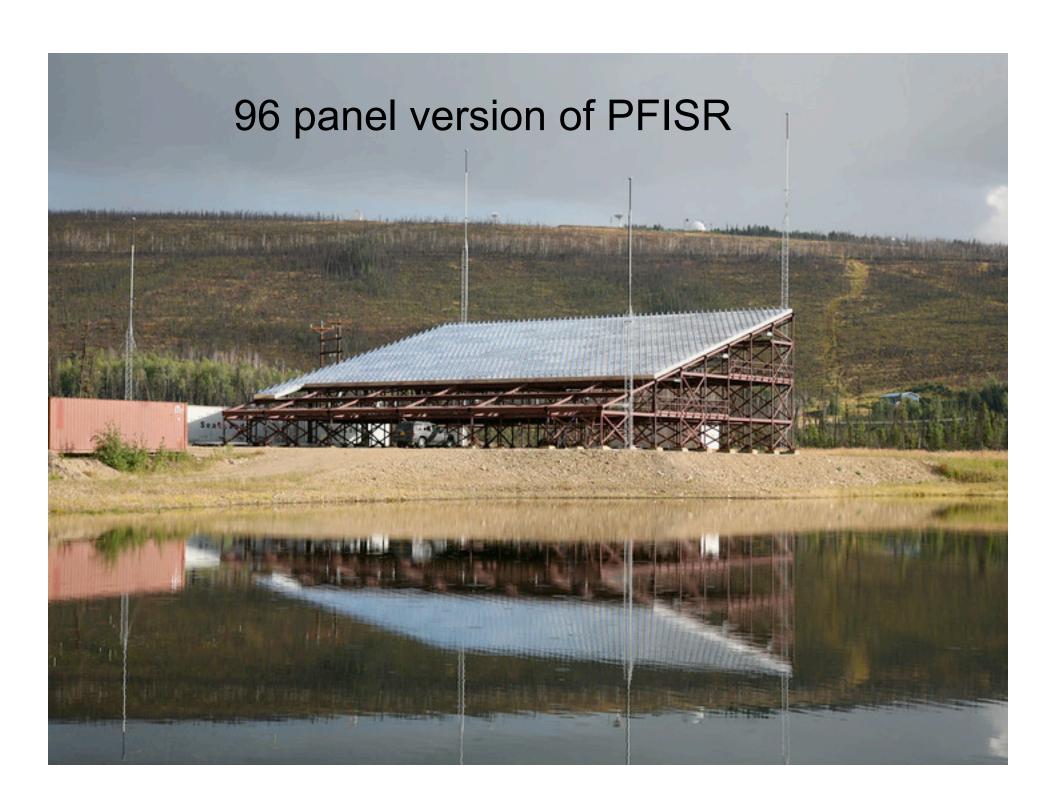


## AMISR Sensitivity vs. Size



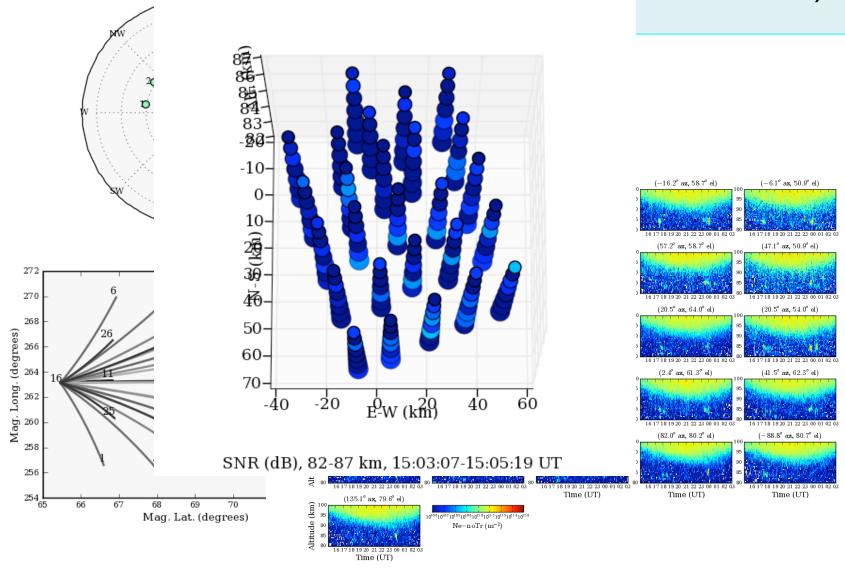
## AMISR Sensitivity vs. Size



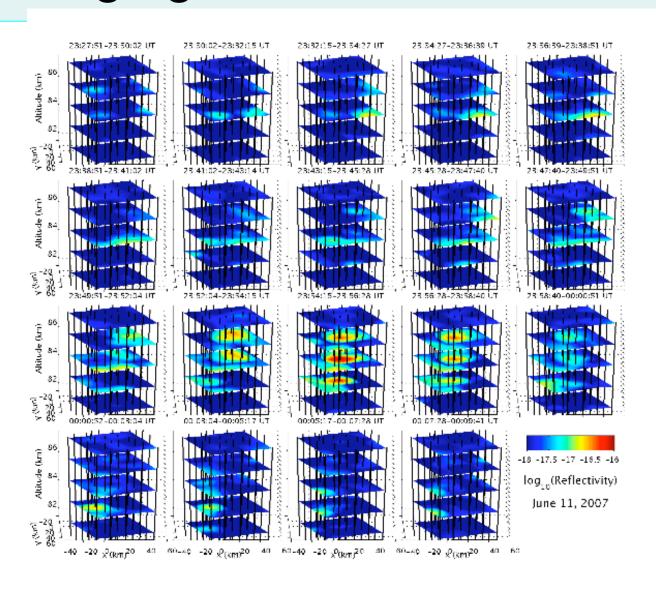




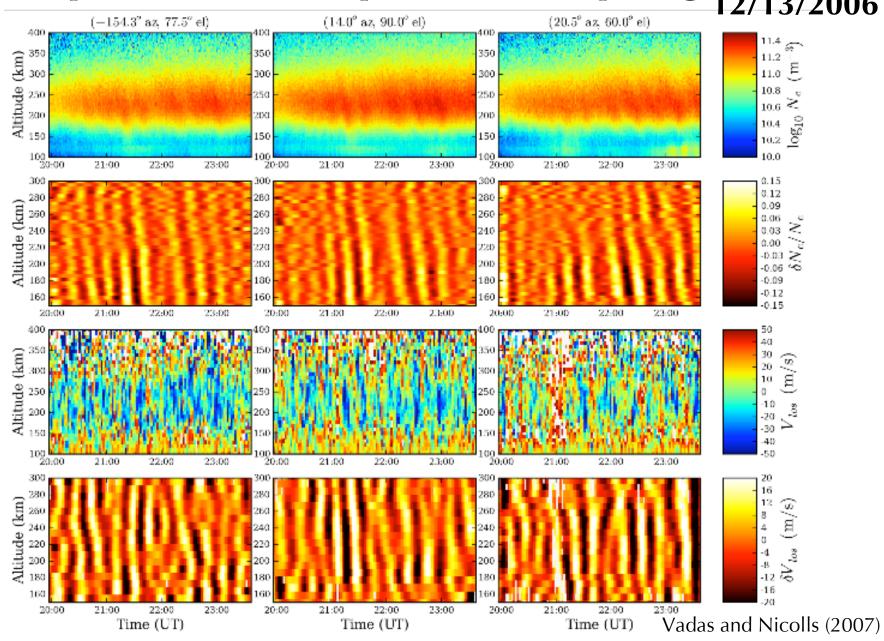
### ospheric Summer Echoes)



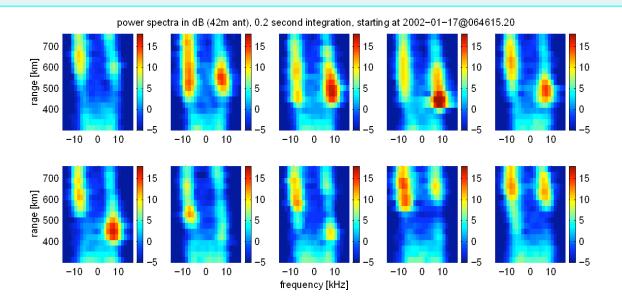
## Imaging PMSE over Poker Flat



## **Ionospheric-Atmospheric Coupling** 12/13/2006

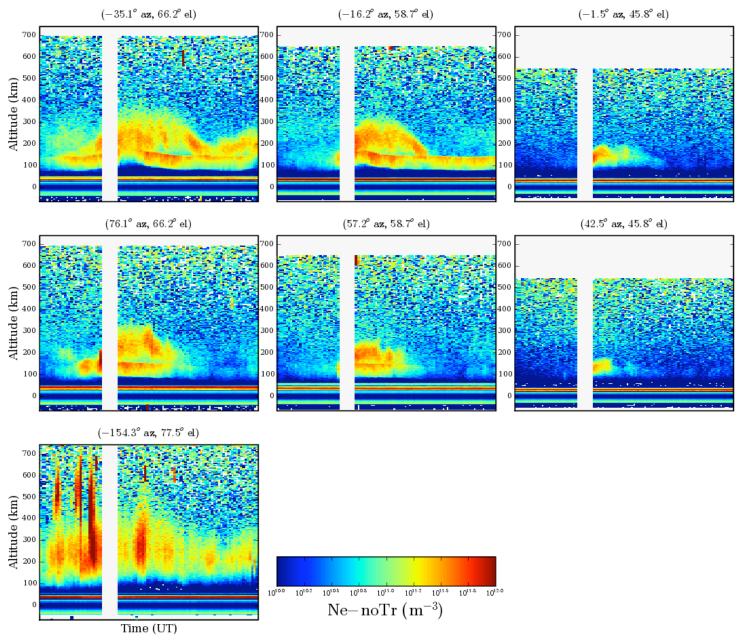


#### Naturally enhanced ion-acoustic lines (NEIAL):

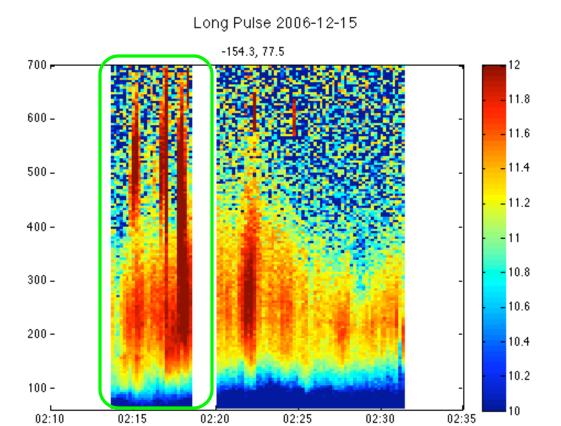


- Order of magnitude enhancement in either or both ion-acoustic lines.
- Over extended altitude range (300 km and up).
- Aligned with the geomagnetic field.
- Rapidly varying in time.

### Raw density in all the 7 positions $_{\rm 12-15-2006~2.228~UT~12-15-2006~2.527~UT}$

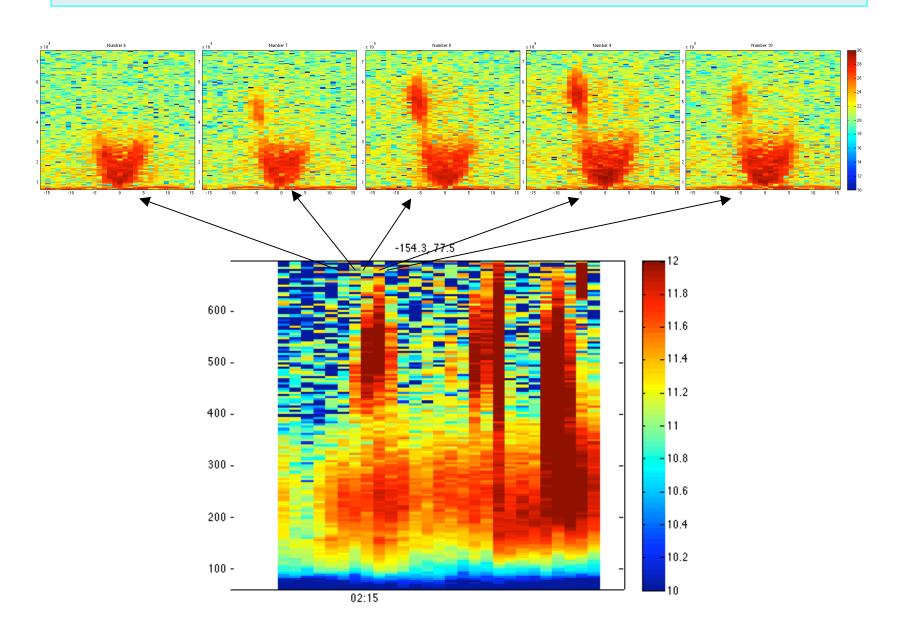


## Strong Coherent Elongated Structures in the upper F-region

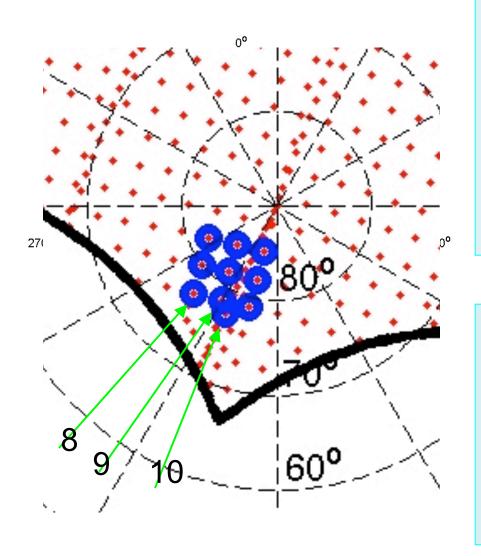


PFISR raw electron densities 15. Dec 2006

### The Ion Line Spectra are Enhanced and Asymmetric



### The Experiment Stromme01

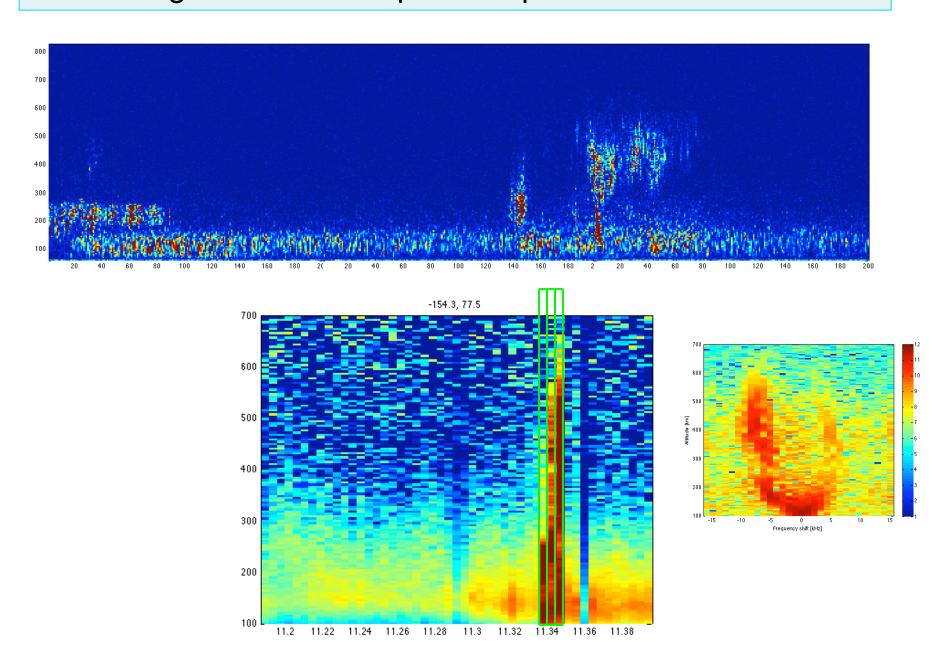


- •9 positions in 3 degree grid
- •10th position up B
- •480 μs pulses (72 km range)
- Raw voltage sampling
- •Plasma line data

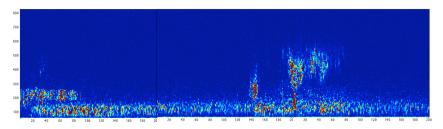
We cycle through the 10 pulses in a fixed order

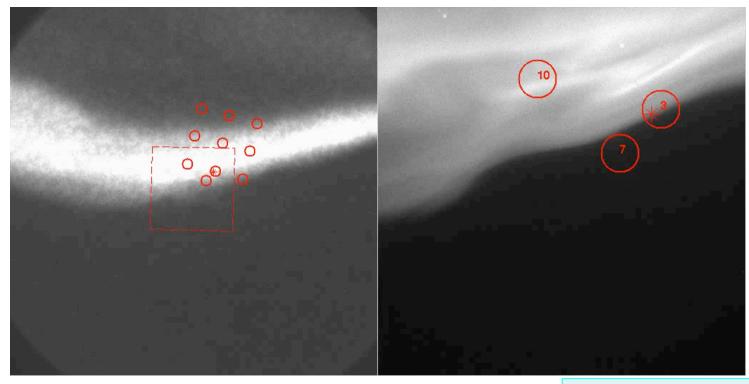
- •7.5 ms between pulses
- •75 ms between pulses in same direction

### Voltage level data -> pulse-to-pulse time resolution



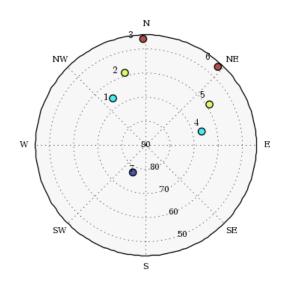
## High Resolution PFISR data and narrow field of view imagers 2007-03-23

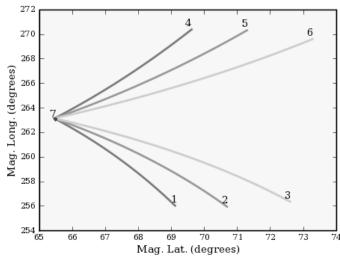




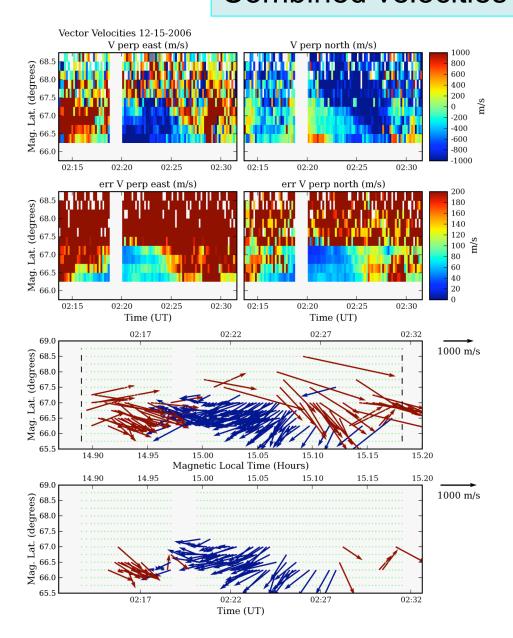
Courtesy of Joshua Semeter

### E-fields with AMISR





### Combined velocities

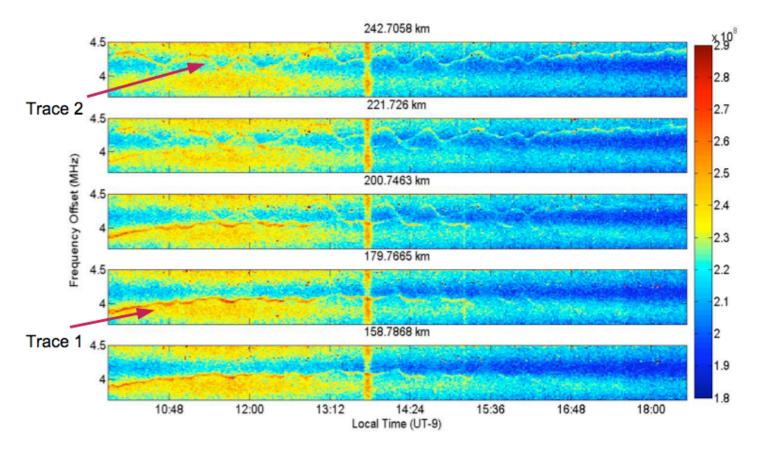


### Plasma lines



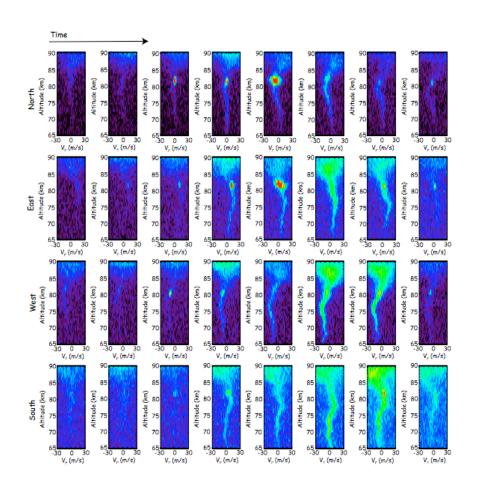


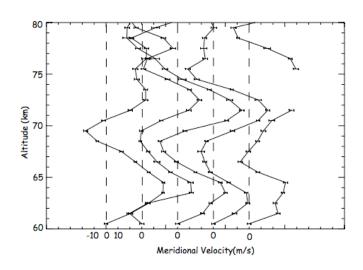
### Plasma line observations



Asti Bhatt 2008

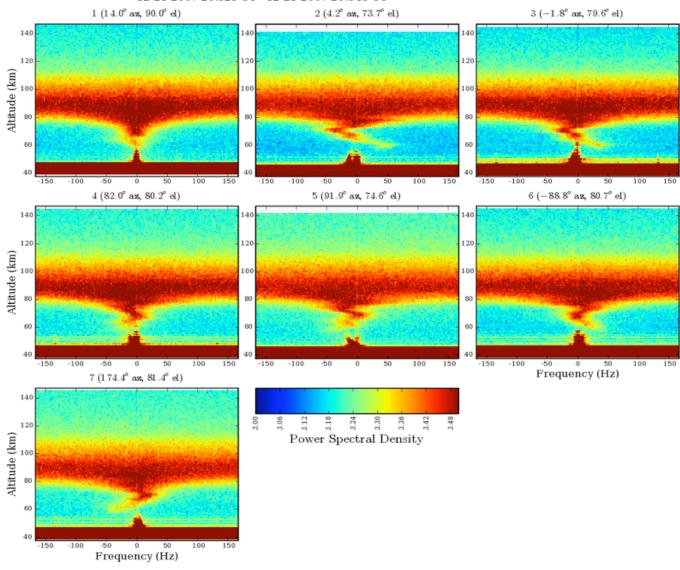
## D region





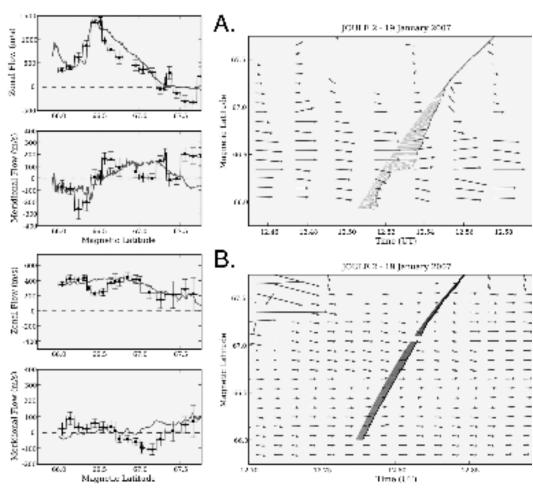
Janches et al 2009

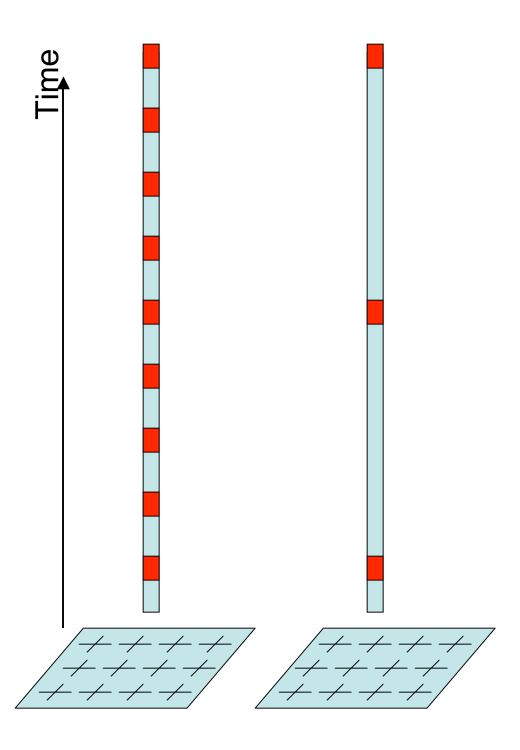
12-21-2007 20.323 UT - 12-21-2007 20.848 UT



### Joule 2 and PFISR

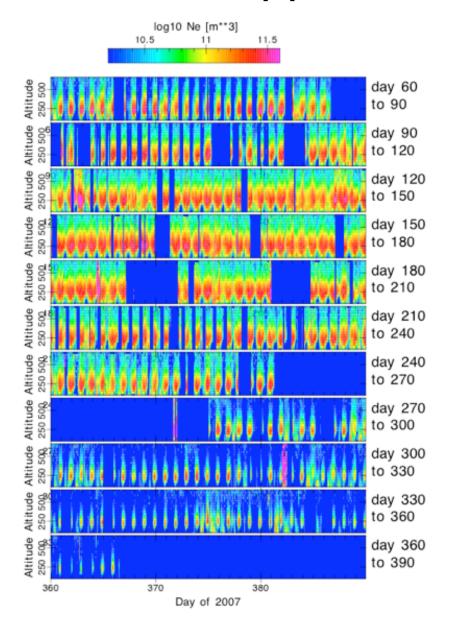




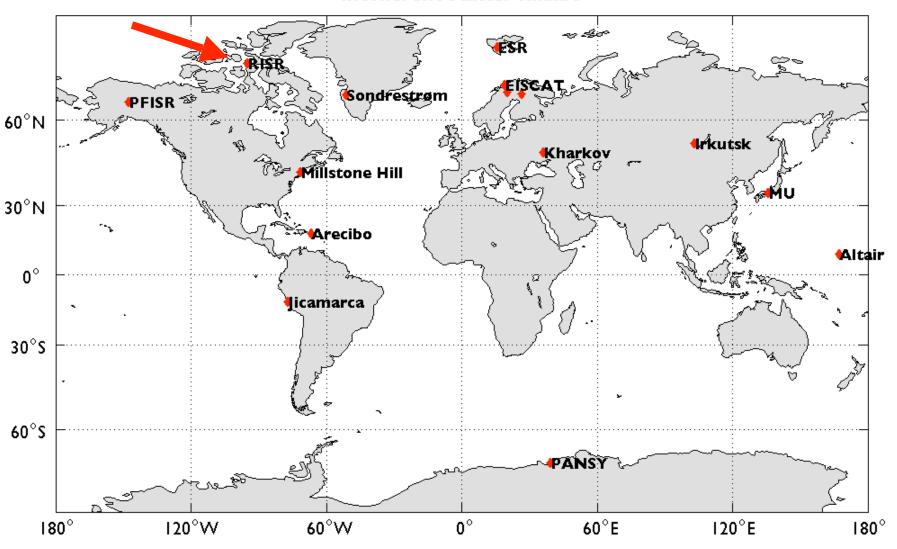


### International Polar Year Support

- EISCAT Svalbard
   Radar and PFISR are
   operating 24 hours per
   day in support of the
   IPY
- Low duty-cycle, single beam mode at PFISR (some augmentation)
- Longest ever IS ionospheric dataset
- Supposed to emphasize "quiet time variability" - coupling from below



#### **Incoherent Scatter Radars**



Map: Thomas Ulich





Photo: John Kelly









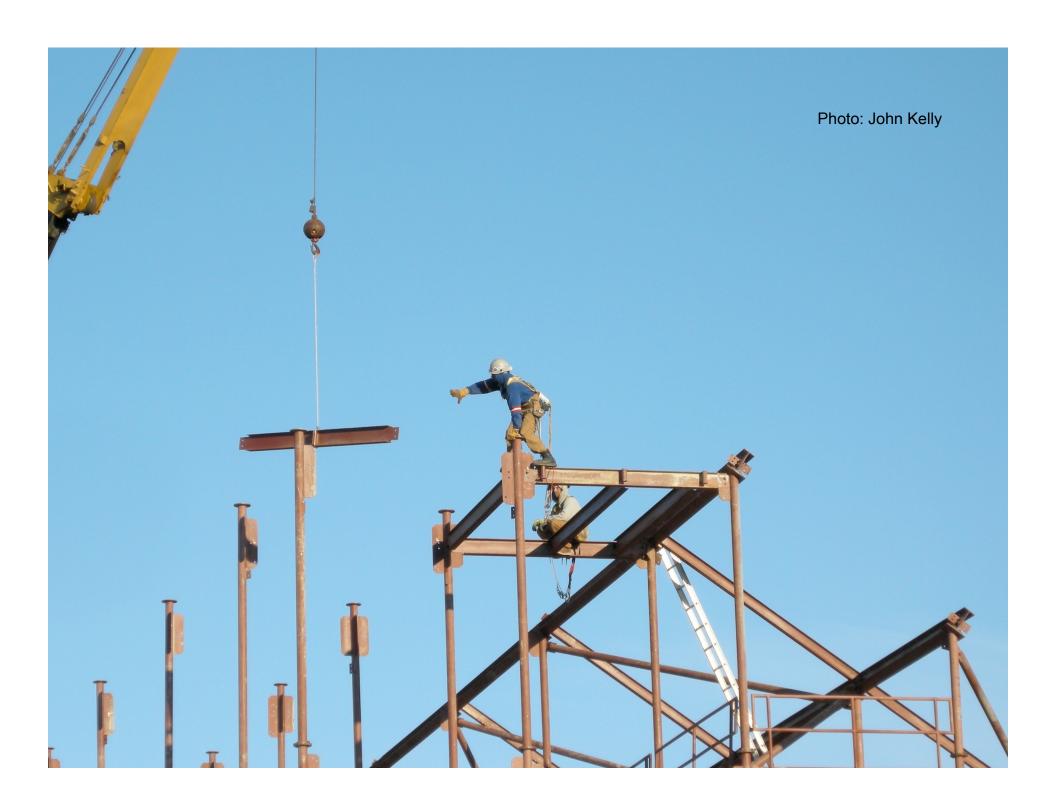


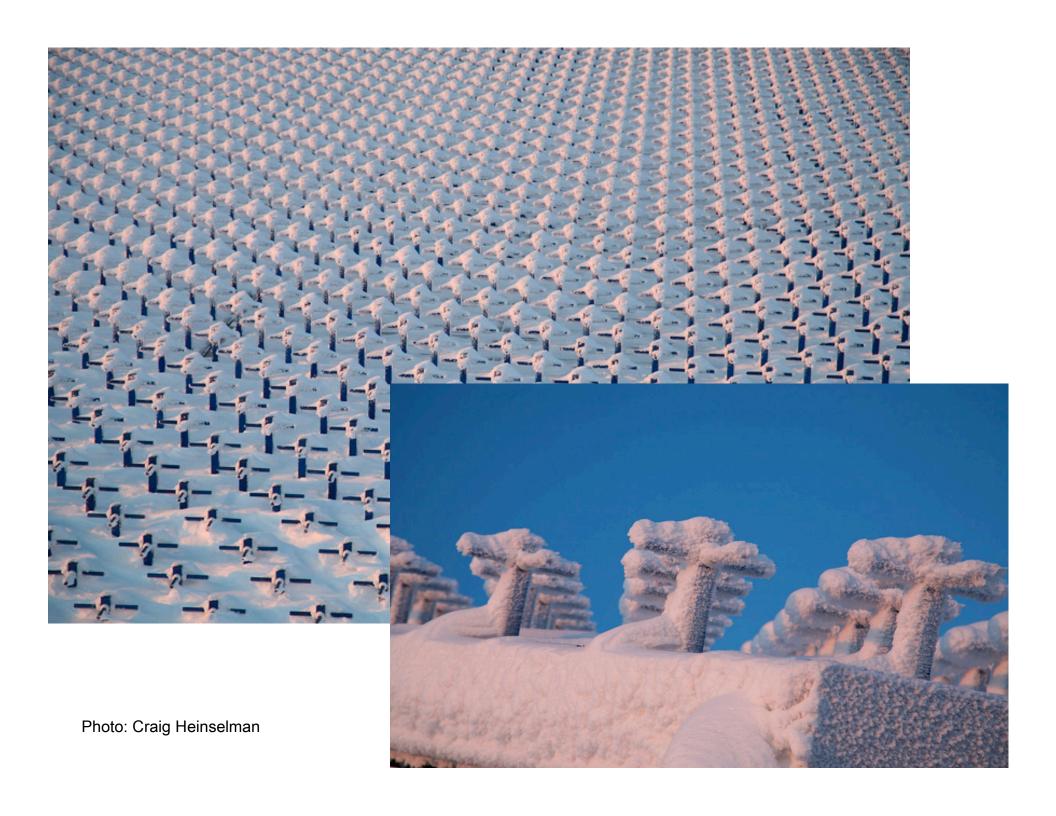




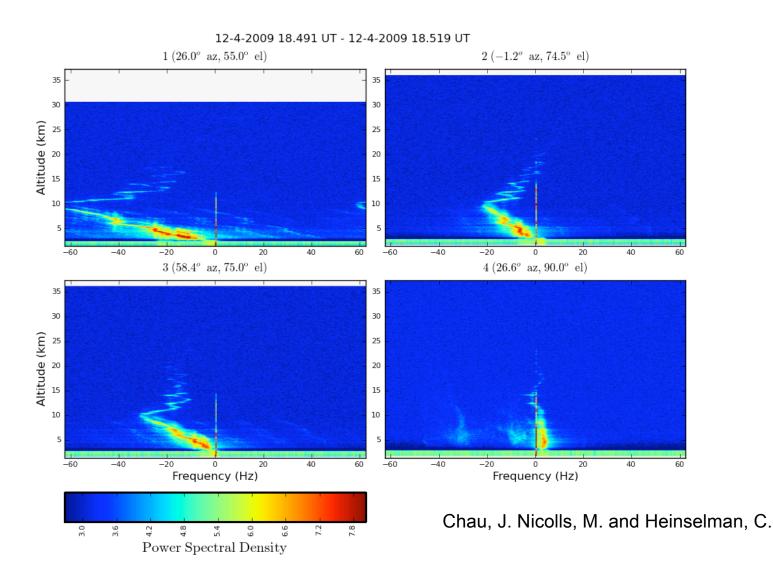


Photo: Craig Heinselman

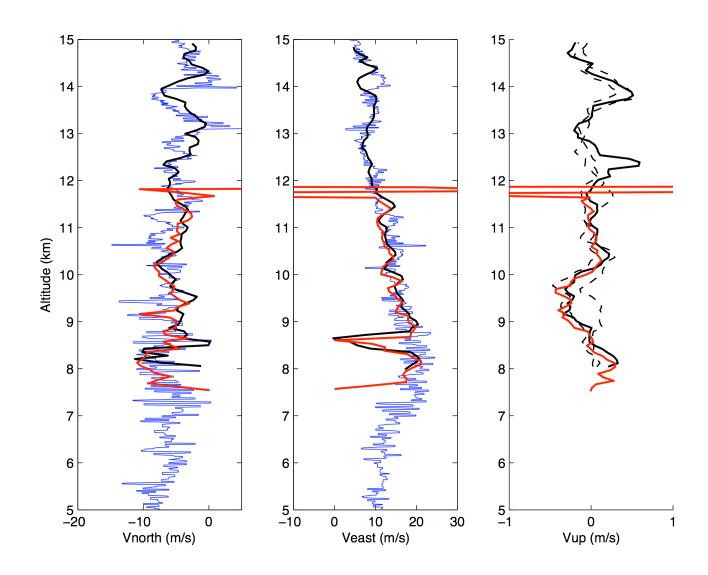




### Pulse to Pulse Spectra at RISR



Resolved winds at PFISR (black and red)
Comparison to balloons (blue)
Note we get vertical velocity – balloon does not



# Other AMISRs...



MUIR - at the HAARP facility in Gakona, Alaska

# **AMISR** at Jicamarca

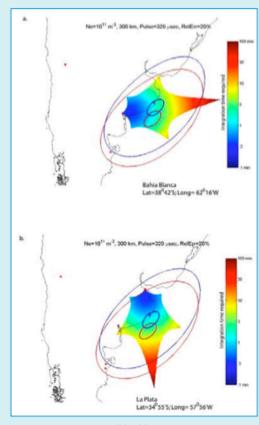


## AMISRs to be...

#### Report on the

Concept Development for an Upper Atmospheric Research Facility at the Arecibo Geomagnetic Conjugate Point in Argentina

NAIC Arecibo Observatory • April 17-19, 2006

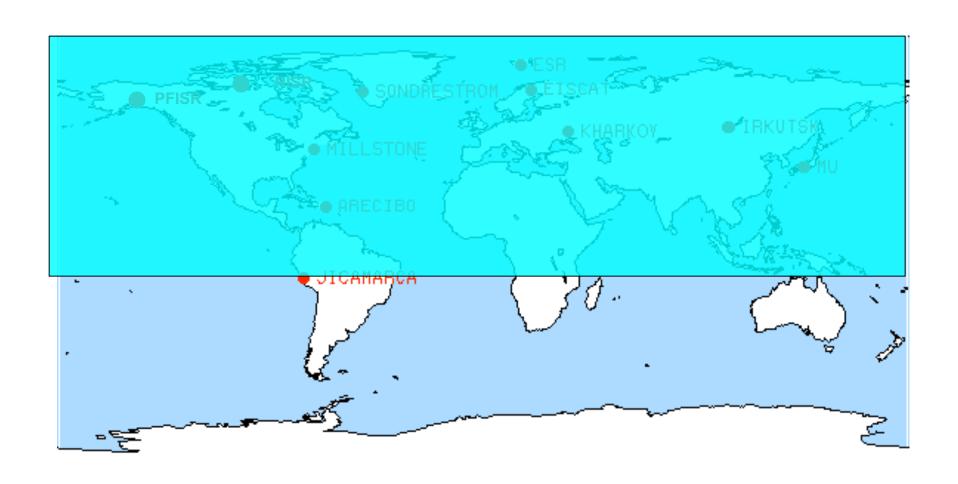


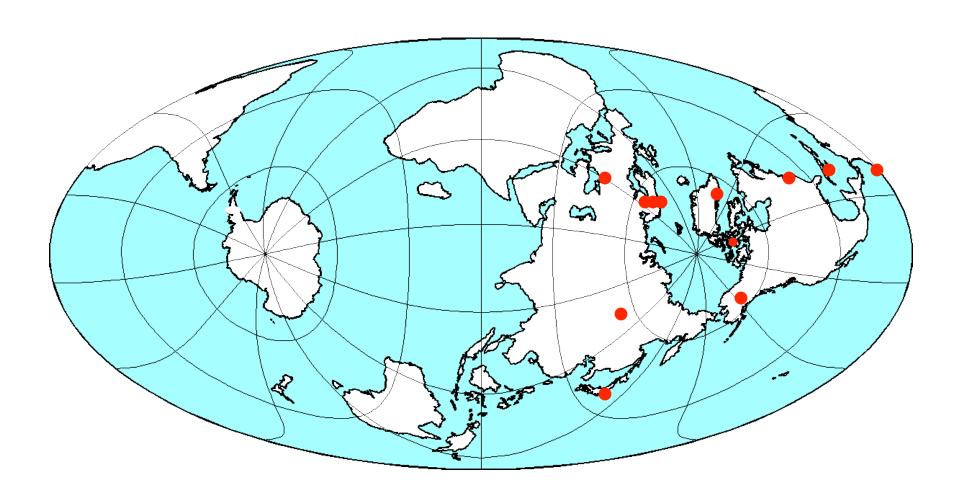
Edited By

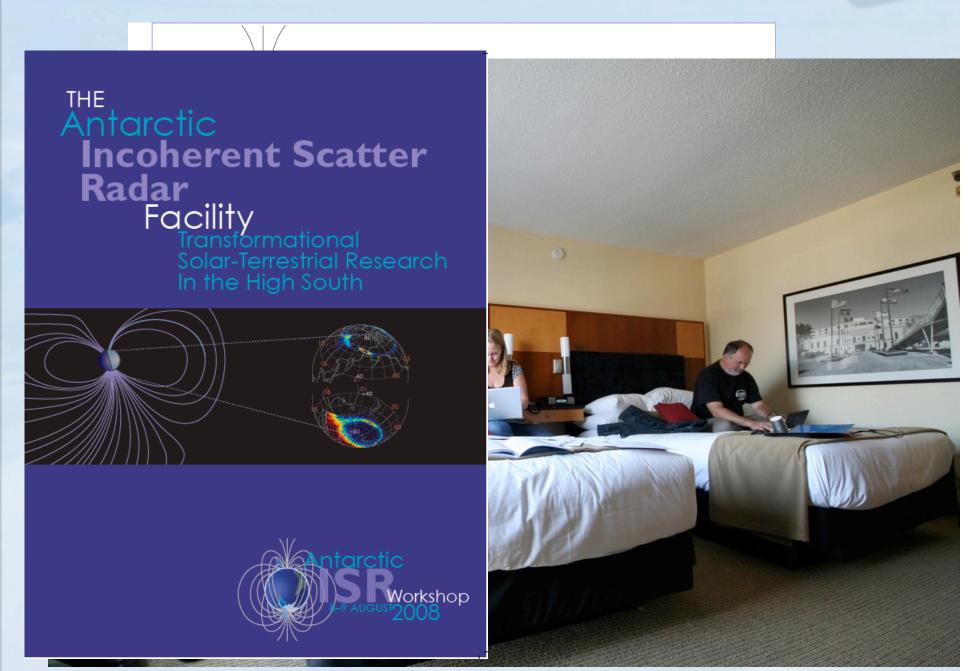
Dr. Diego Janches, Northwest Research Associates

Dr. Robert L. Brown, National Astronomy and Ionosphere Center

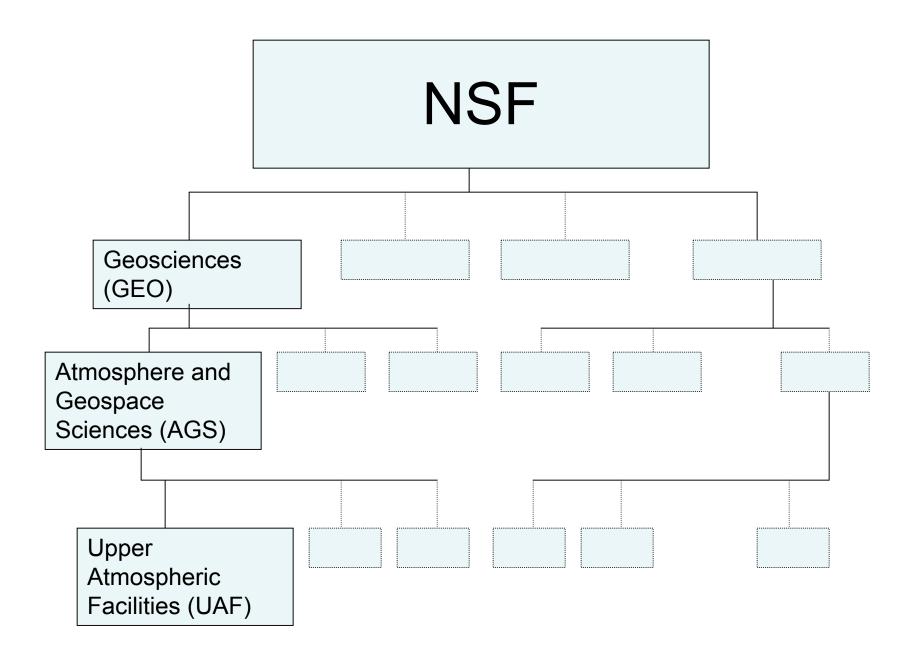
### Incoherent Scatter Radars of the World

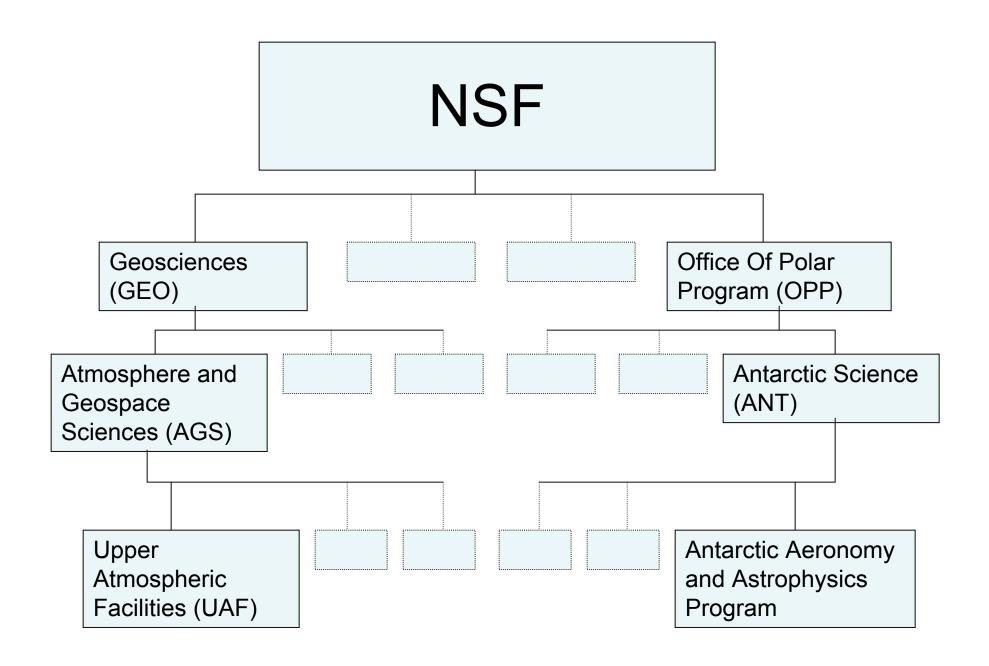




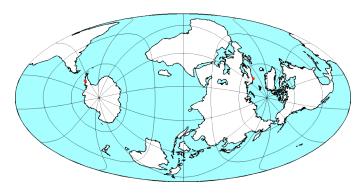


ftp://isr.sri.com/pub/Antarctic-ISR/Antarctic\_ISR\_Workshop\_2008.pdf

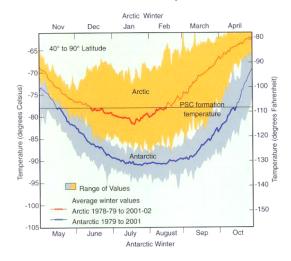


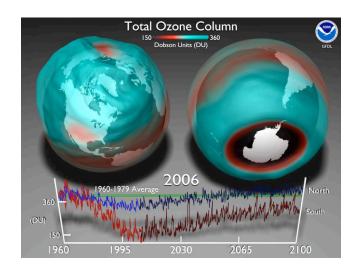


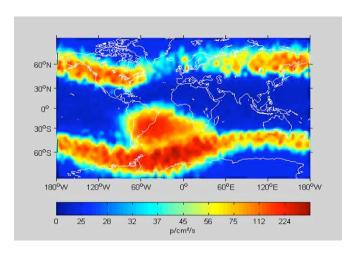
## Antarctica is different



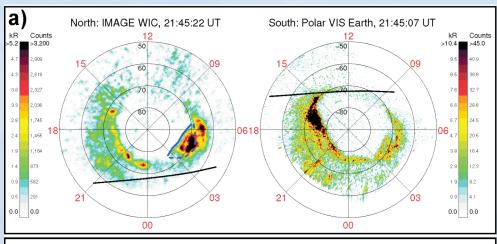
#### Minimum Air Temperatures in the Polar Lower Stratosphere

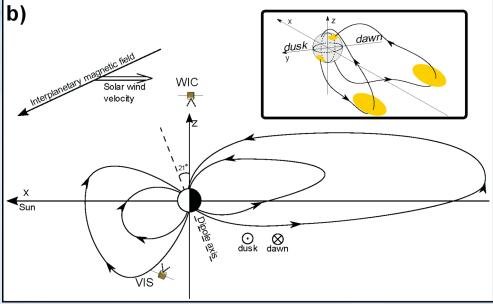






### Antarctica is different





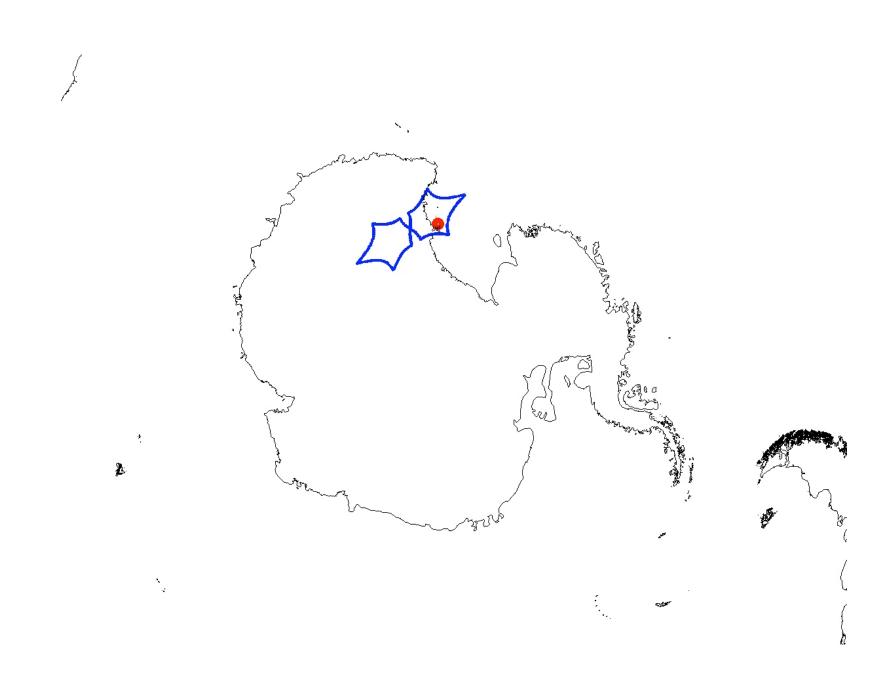


From Laundal and Østgaard 2009

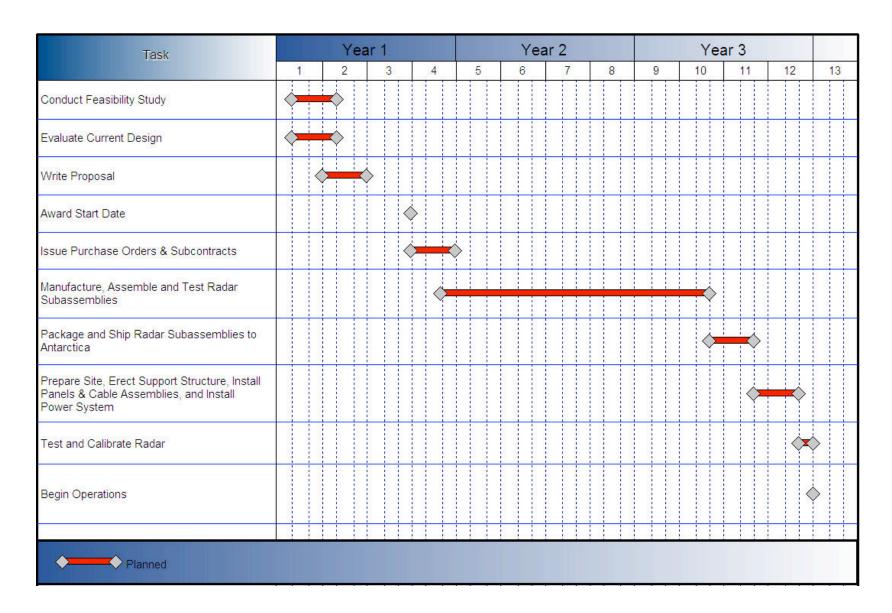
# Antarctic Stations and the conjunct points for the northern ISRs Rothera Halley Sanae McMurdc Sondrestrom Poker • Resolute Vostok Showa Davis Shan • EISCAT

# Some possible locations

Conjugate pair	Geographic lat & lon		Geomagnetic lat & lon	
Syowa	-69.0,	39.6	-70.4,	83.6
Leirvogur (Iceland)	64.2,	338.3	69.3,	71.1
McMurdo	-77.9,	166.7	-79.0,	290.1
Resolute	74.7,	265.1	82.9,	303.0
Davis	-68.6,	78.0	-76.5,	128.2
Zhonshan	-69.4,	76.4	-74.6,	96.5
Longyeabyen	78.2,	15.8	75.3,	112.1

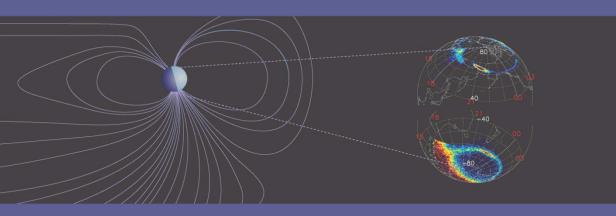


## Timeline:









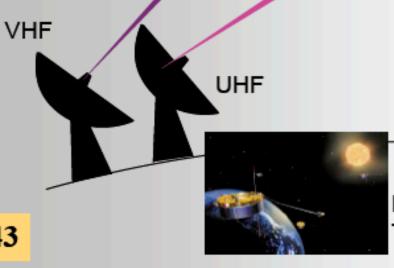
ftp://isr.sri.com/pub/Antarctic-ISR/Antarctic\_ISR\_Workshop\_2008.pdf





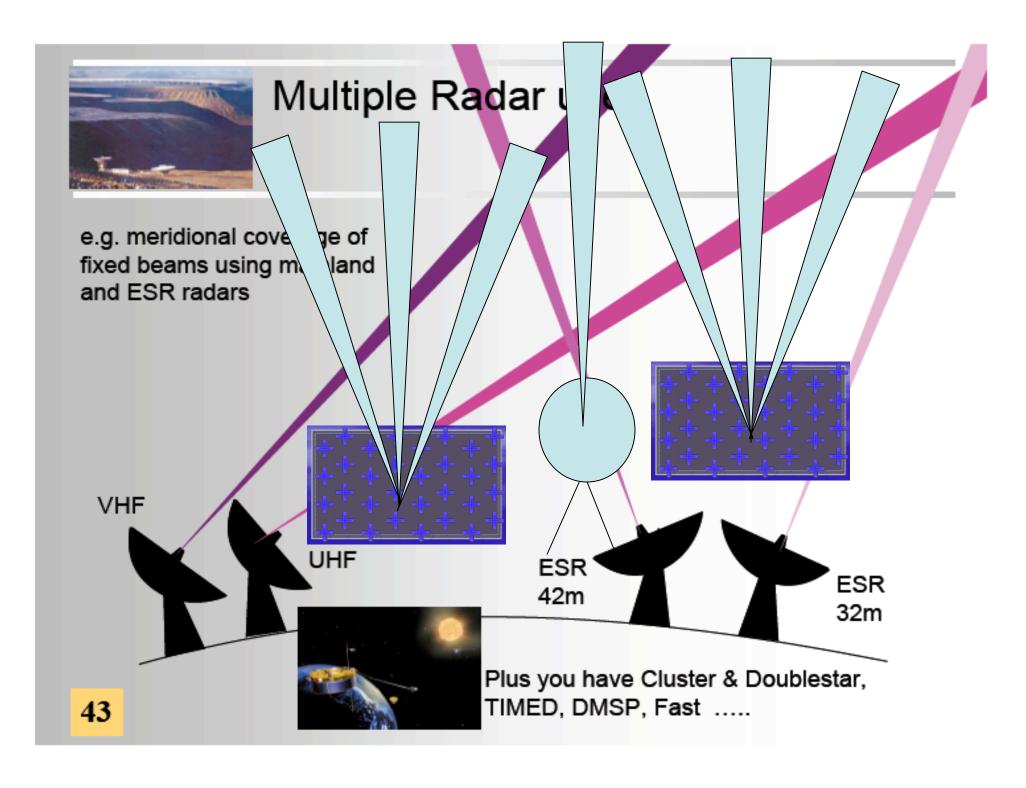
## Multiple Radar use

e.g. meridional coverage of fixed beams using mainland and ESR radars



**ESR ESR** 42m 32m

Plus you have Cluster & Doublestar, TIMED, DMSP, Fast .....



# "Summary"

- As scientists we should focus more on the science and processes then on the instruments
- There are no (national or institutional) borders in space!

